

Intensity of Bragg scattering is proportional to
 (a) e^{-2W}

where $W = \frac{1}{N} \sum_{\vec{k}, s} \frac{\hbar}{2M\omega_{s,\vec{k}}} (\vec{Q} - \vec{\epsilon}_{s,\vec{k}})^2 \left[n(\omega_{s,\vec{k}}) + \frac{1}{2} \right]$

Averaging over directions, which gives an unimportant numerical factor of order unity, and replacing the sum by an integral gives

$$W \propto \int_0^\infty dk k^{d-1} \frac{1}{\omega_k} \left[\frac{1}{e^{sk\omega_k} - 1} + \frac{1}{2} \right]$$

Now $\omega_k \propto k$ and we are interested in the ^{poor} divergent behavior of the integrand as $k \rightarrow 0$, so

~~1/2~~
 (8) $W \propto \int_0^\infty \frac{dk}{k} k^{d-1} \left[\frac{T}{k} + \frac{1}{2} \right]$ neglecting constants

For $T > 0$ the T/k term dominates over $1/2$

$$W \propto T \int_0^\infty \frac{dk}{k} k^{d-2}$$

In $d=2$ $W \propto T \int_0^\infty \frac{dk}{k}$ which diverges logarithmically

In $d=1$ $W \propto T \int_0^\infty \frac{dk}{k^2}$ " diverges linearly

A.B. in $d=3$ W is finite.

(b) $T=0$ $W \propto \int_0^\infty \frac{dk}{k} k^{d-1}$

In $d=2$ $W \propto \int_0^\infty dk \text{const.}$ which is finite.

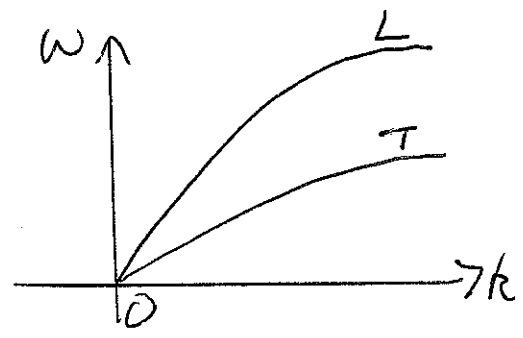
In $d=1$ $W \propto \int_0^\infty \frac{dk}{k}$ which diverges logarithmically.

(c) If $W = \infty$, $e^{-2W} = 0$ so there is no translational long range order when W diverges (though one has "quasi long range order" in 2-d (at finite $-T$), see the hand out.

3/ Dispersion relation

$$\frac{d^2\omega}{dk^2} < 0$$

3 phonon processes

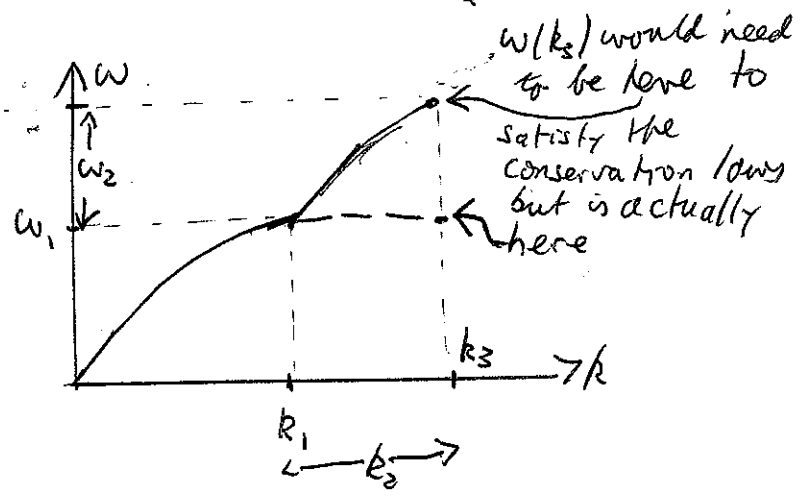


$$\omega_1 + \omega_2 = \omega_3$$

$$k_1 + k_2 = k_3 \quad (\neq G)$$

(a) all phonons belong to the same branch.

Graphical construction \Rightarrow



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(b) Since $\omega_L(k) > \omega_T(k)$ the process

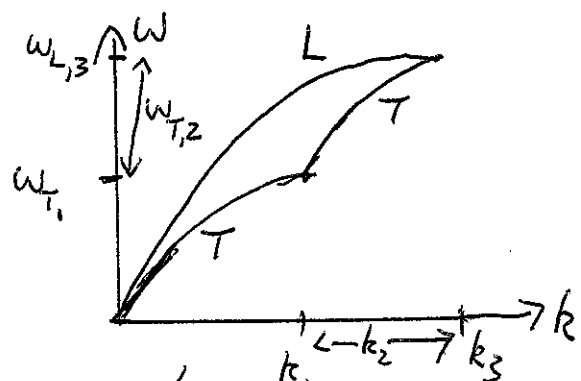
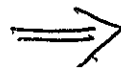
$$L + L \leftrightarrow T$$

also does not work, because the mismatch in frequencies (if k conservation is satisfied) is even greater.

Similarly for $L + T \leftrightarrow T$.

In order to make up the mismatch in frequencies the single phonon must be in a branch higher than at least one of the members of the pair, eg

$T + \tau \leftrightarrow L$

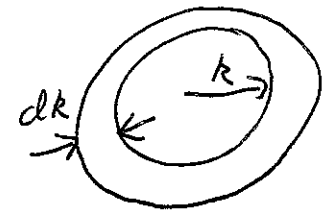


Similarly $T + L \leftrightarrow L$ works

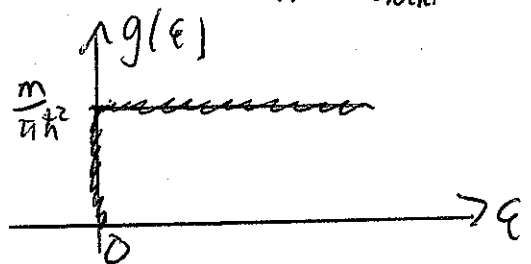
3/ (a) Density of points in k -space / unit volume, is $\frac{1}{(2\pi)^2}$
 Hence $n = \frac{2 \uparrow \text{spin}}{\pi k_F^2} \frac{1}{(2\pi)^2} = \frac{k_F^2}{2\pi}$
 πk_F^2 is the area of Fermi circle

(b) $n = \frac{1}{\pi r_s^2}$ πr_s^2 is the area per particle.
 i.e. $\frac{1}{\pi r_s^2} = \frac{k_F^2}{2\pi}$ or $k_F = \frac{\sqrt{2}}{r_s}$

(c) $g(\epsilon) d\epsilon = \frac{2 \uparrow \text{spin}}{2\pi k} \frac{dk}{\text{area of shell}} \frac{1}{(2\pi)^2}$



(8) $g(\epsilon) = \frac{k}{\pi} \frac{1}{d\epsilon/dk} = \frac{k m}{\pi \hbar^2 k} = \frac{m}{\pi \hbar^2}$, a constant (for $\epsilon > 0$)



(a) Sommerfeld expansion

$$n = \int_0^{\infty} f(\epsilon) g(\epsilon) d\epsilon = \int_0^{\mu} g(\epsilon) d\epsilon + \sum_{n=1}^{\infty} a_n (k_B T)^{2n} \frac{d^{2n-1}}{d\epsilon^{2n-1}} g(\epsilon) \Big|_{\mu}$$

But all derivatives vanish so

$n = \int_0^{\mu} g(\epsilon) d\epsilon$ according to the Sommerfeld expansion

At $T=0$, $n = \int_0^{\epsilon_F} g(\epsilon) d\epsilon$ so $\mu(T) = \epsilon_F$ according to Somm. expansion.

(e) Does this mean that $\mu(T)$ is really identical to E_F ($\equiv \mu(T=0)$)?

NO. It just means that all terms that are powers of T vanish. As we shall see in ~~this part~~ and the next section there are corrections which vanish exponentially as $T \rightarrow 0$.

Now $n = \frac{m}{\pi k^2} \int_0^{E_F} d\epsilon = \frac{m E_F}{\pi k^2}$ (1) ($T=0$)

and $n = \frac{m}{\pi k^2} \int_0^\infty \frac{1}{e^{\beta(\epsilon-\mu)} + 1} d\epsilon = \frac{m k_B T}{\pi k^2} \int_{-\beta\mu}^\infty \frac{dx}{e^{x+1}} \quad (T > 0)$

Now $I = \int \frac{dx}{e^{x+1}} = \int \left[\frac{1}{e^{x+1}} \right] dx = x - \ln(1+e^x)$ so

$n = \frac{m k_B T}{\pi k^2} \left[\frac{\mu}{k_B T} + \ln(1 + e^{-\mu/k_B T}) \right]$ (2)

Equate (1) and (2) $\ln[1 + e^{-\mu/k_B T}]$, so from (1) and (2) $e^{\beta\mu} = e^{\beta E_F} + 1$
 $E_F = \mu + k_B T \ln[1 + e^{-\mu/k_B T}]$ (3) can also be written
 $\mu = k_B T \ln[e^{\beta E_F} - 1]$ (3)

(f) As $T \rightarrow 0$, μ is very close to E_F . Also $e^{-\mu/k_B T}$ is very small and so, to leading order in small quantities, we can replace it by $e^{-E_F/k_B T}$.
Hence $\mu(T) \approx E_F - k_B T \ln[1 + e^{-E_F/k_B T}]$
 $\approx E_F - k_B T e^{-E_F/k_B T} \dots$

i.e. $\mu - E_F$ is exponentially small as $T \rightarrow 0$

(g) (not asked this year) at high T , from (3) $\mu = k_B T \ln(\beta E_F) = k_B T \ln\left(\frac{n \pi k^2}{m k_B T}\right) \quad (< 0)$
 $= k_B T \ln(n \lambda_Q)$ $\lambda_Q = \frac{\pi k^2}{m k_B T}$ is the square of the thermal de Broglie wavelength

4/ $e^{\mu/k_B T} \ll 1$ for classical statistics

Use the classical expression

(a) $n = \int_0^\infty e^{-(\epsilon-\mu)/k_B T} g(\epsilon) d\epsilon.$

Now $g(\epsilon) = \frac{\sqrt{2}}{\pi^2} \left(\frac{m}{\hbar^2}\right)^{3/2} \epsilon^{1/2}$ so

$$n = \frac{\sqrt{2}}{\pi^2} \left(\frac{m}{\hbar^2}\right)^{3/2} e^{\beta\mu} \int_0^\infty \epsilon^{1/2} e^{-\beta\epsilon} d\epsilon.$$

i.e. $n = \frac{1}{\sqrt{2}} \left(\frac{m \pi \hbar^2}{\hbar^2 k_B T}\right)^{3/2} e^{\mu/k_B T}.$ $\uparrow (3/2) = \frac{1}{2} \sqrt{\pi}$

But this is only valid if $e^{\mu/k_B T} \ll 1$

i.e. $n \ll \frac{1}{\sqrt{2}} \left(\frac{m \pi \hbar^2}{\hbar^2 k_B T}\right)^{3/2}.$

(6) But $n = \frac{1}{\frac{4\pi}{3} \tau_s^3}$, τ_s is the typical inter particle spacing

i.e. $\tau_s \gg \left(\frac{\hbar^2}{2mk_B T}\right)^{1/2}$ (dropping factors of order unity)

(b) $\left(\frac{\hbar^2}{2mk_B T}\right)^{1/2}$ is the thermal de-Broglie wavelength, i.e. the de Broglie wavelength of a particle whose energy is $k_B T$ (times $\frac{1}{2\pi}$)

i.e. $\frac{\hbar^2}{2m} \left(\frac{2\pi}{\lambda_{deB}}\right)^2 = k_B T$

$\left(k_{deB} = \frac{1}{\lambda_{deB}}\right)$

(c) ~~just plug in the numbers~~ See next page

5 $V_F = \frac{\hbar}{m} k_F$

(a) Roughly, $k_F \sim \frac{1}{a_0}$ where a_0 is the Bohr radius (i.e. of order the inter-atomic separation)

$a_0 = \frac{\hbar^2}{me^2}$

Hence $V_F \simeq \frac{\hbar}{m} \frac{me^2}{\hbar^2} = \frac{e^2}{\hbar} = \frac{e^2}{\hbar c} c = \alpha c$

(3) where $\alpha = \frac{e^2}{\hbar c} \simeq \frac{1}{137}$ is the fine structure constant

(b) The ions are roughly ⁴⁶⁵⁵ orders of magnitude heavier than the electrons and, in general, $\omega \propto M^{-1/2}$ (see e.g. the simple harmonic oscillator)

4 (c) $\hbar = 1.05 \times 10^{-34}$, $m = 9.1 \times 10^{-31}$, $k_B = 1.38 \times 10^{-23}$

$a_0 = 0.53 \times 10^{-10}$
condition (1) on p. 5 is
Hence $\sqrt{\frac{\hbar^2}{2mk_B T a_0^2}}$ \Rightarrow $\left[\frac{(1.05 \times 10^{-34})^2}{2 \times 9.1 \times 10^{-31} \times 1.38 \times 10^{-23} \times (0.53 \times 10^{-10})^2} \right]^{1/2}$
 $= \left[\frac{1.6 \times 10^5}{T} \right]^{1/2} \simeq \left[\frac{10^5}{T} \right]^{1/2}$